

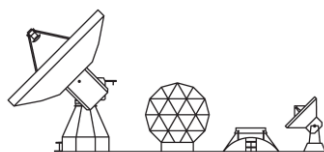
QRFH antenna for FGI VGOS receiver (YQR-01-001)

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Abstract– This report summarizes the Quad-Ridged Feed Horn (QRFH) results for Yebes built model YQR-01-001. It includes preliminary specifications, and interface drawings and experimental results of the manufactured antenna. Reflection, isolation, and radiation patterns have been measured in the 2.2 to 14 GHz frequency range. In addition, aperture efficiency on a VGOS 13.2m ring-focus radio telescope has been estimated by numerical calculations using the measured radiation patterns as input.



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1 Introduction

This document presents the development of an antenna feed to be part of a VGOS broadband receiver that it is being built at Yebes Observatory for the Finnish Geospatial Research Institute (FGI) new radio telescope located at the Metsähovi Geodetic Fundamental Station. The feed follows a broadband dual-polarized QRFH topology, which is being widely utilized in many broadband radio astronomy systems [1][2].

The electrical design has been carried out by optimizing the horn and ridge's profiles. The goals for the optimization were input reflection losses under a given threshold, and aperture efficiency on the reflector as high as possible. Two SMA compatible connectors are used to get both linear orthogonal polarizations.

The antenna has been manufactured by the Spanish industry, whereas the design and tests have been carried out at Yebes Observatory. The antenna description and specifications, as well as the measured parameters are summarized in Table I.

Project	VGOS [4]	
Application	Antenna Feed for VGOS 13.2m radio telescope	
Where to	Finnish Geospatial Research Institute (FGI)	
Manufactured by	UTILMEC, S.L. (Spain)	
Date	2015	
Output	2 coaxial SMA type	
Reference impedance	50 Ω	
Frequency	2.2-14 GHz	
	Specification	Measured
Port reflection(typ.) (max.)	-8 dB	-8 dB -5 dB
Port isolation (min.)	20 dB	20 dB
Mean aperture efficiency over VGOS radio telescope (calculated from measurements)	50 %	60 %
Xpolar (min.)	20 dB	20 dB

Table I. Summary of specifications and results of the manufactured QRFH antenna.

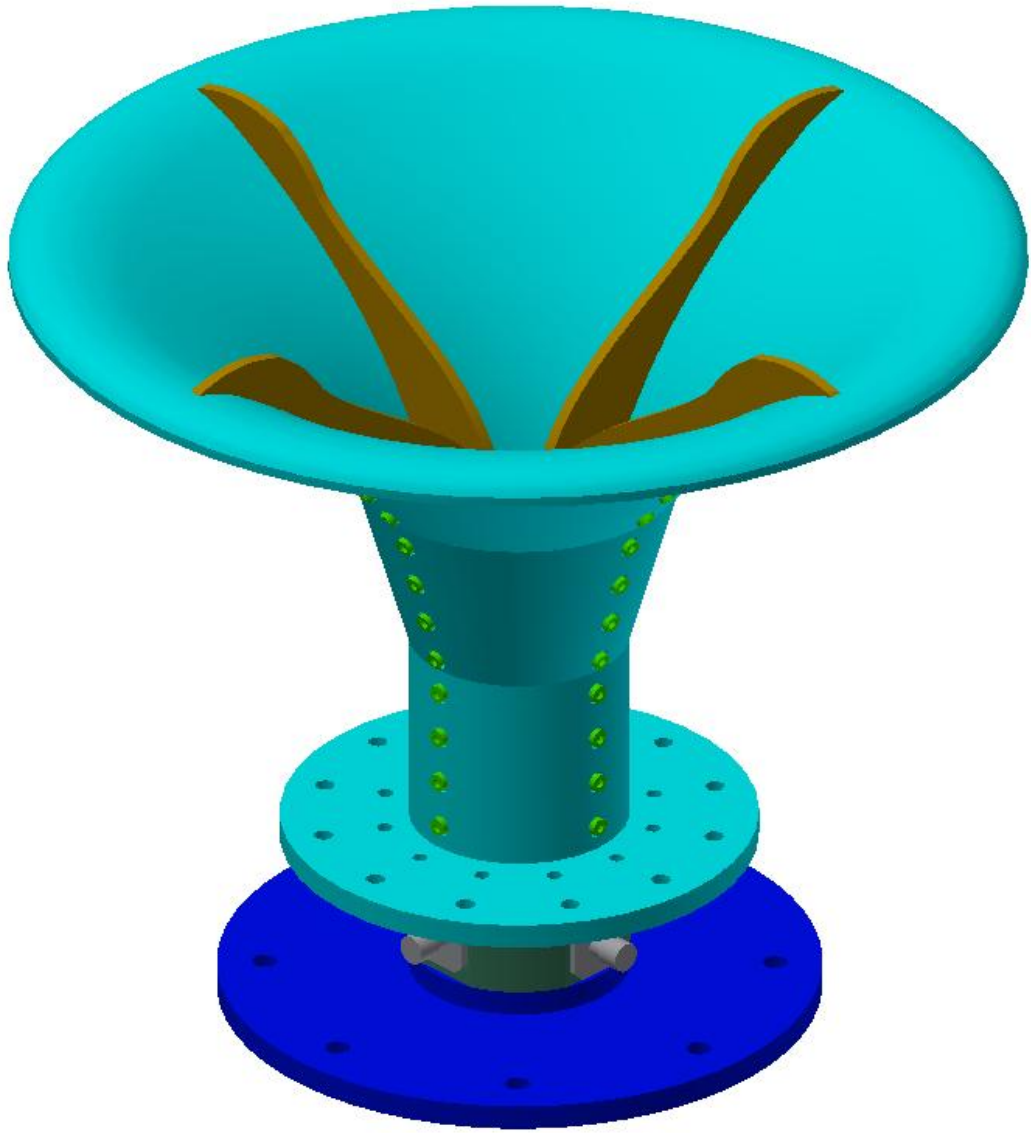


Fig 1.1. 3D Model of QRFH.

2 Mechanical interface

The antenna consists of an external horn and four internal ridges (QRFH [2]), whose profiles are defined by analytical curves. The ridges slightly protrude above the horn at the end of the aperture. However, for convenience, the aperture has been defined as the edge of the horn profile. The antenna also has a transition between the quad-ridged waveguide and the two coaxial ports. At the bottom of the antenna there is a cover with the appropriate interface to attach the antenna at the cold stage of the cryostat. The antenna is entirely made of aluminum.

Fig 2.1 shows the profile view of the antenna, including the distances from the bottom to the limit of the ridges, the aperture, the phase center (the calculation of its position is described in Section 3.3) and the two coaxial ports respectively. The interface view at the bottom of the antenna is shown in Fig 2.2. Ports position from the bottom view is shown in Fig 2.3.

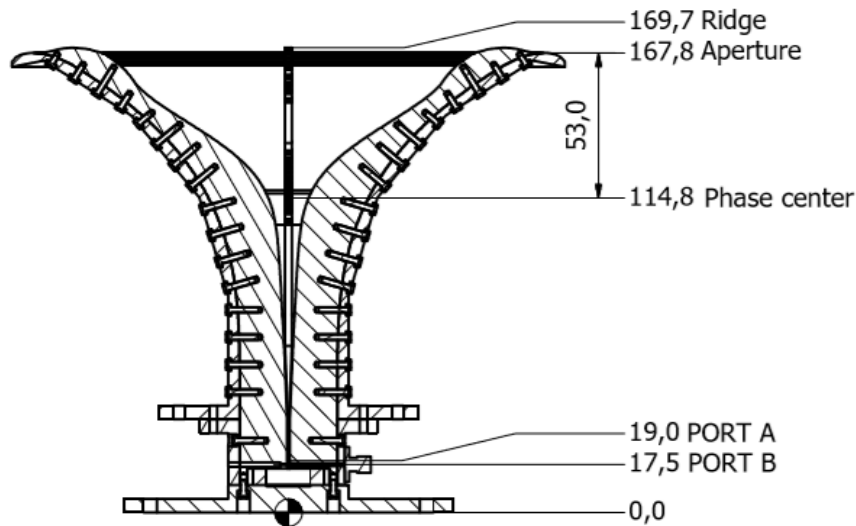


Fig 2.1. Profile view of the QRFH.

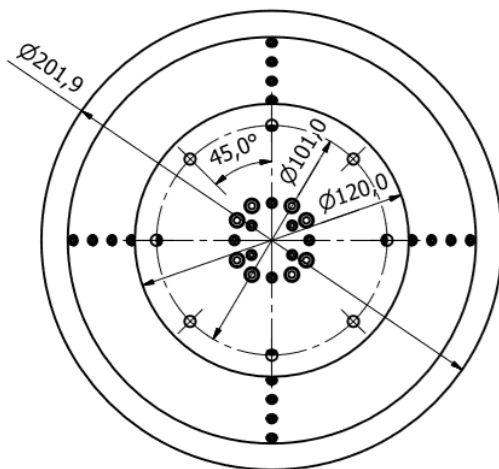


Fig 2.2. Interface view of the antenna.

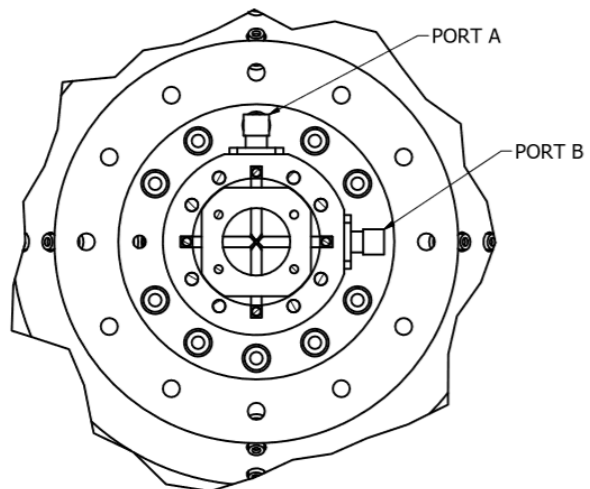


Fig 2.3. Ports position from bottom view without cover.

3 Feed measurements

3.1 Port reflection and isolation

The return losses have been measured at the output coaxial connectors, obtaining the results shown in Fig 3.1. On the other hand, the isolation between both ports is shown in Fig 3.2.

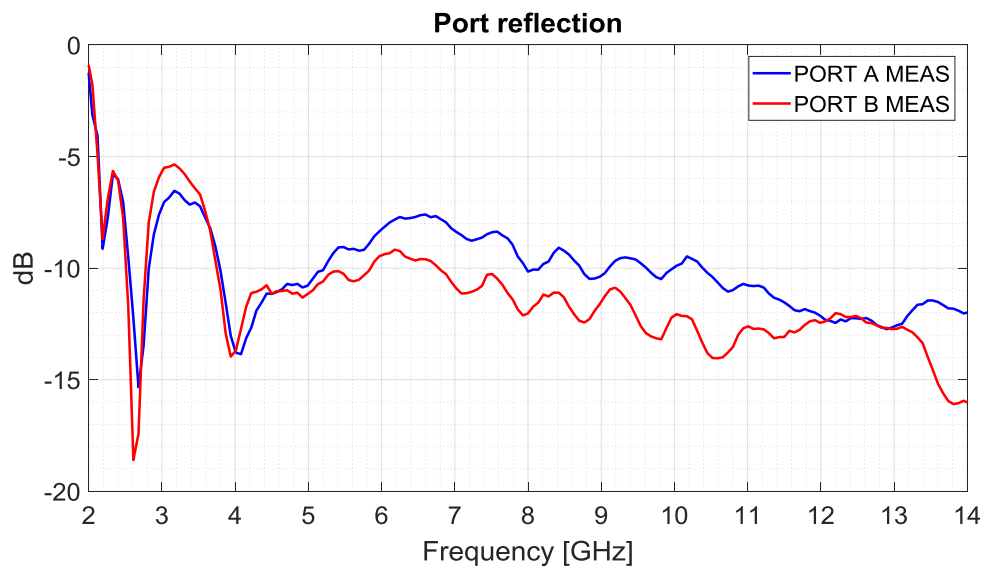


Fig 3.1. Measured port reflection

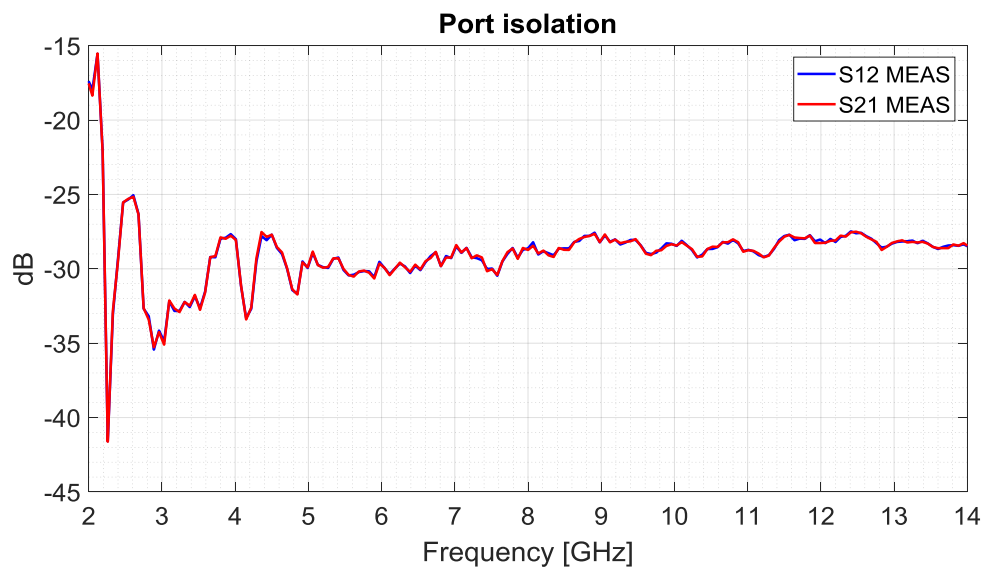


Fig 3.2. Measured port isolation.

3.2 Radiation patterns

The measurement of the antenna radiation patterns has been performed at the anechoic chamber of Yebes (CDTAC) (Fig 3.3)[5].

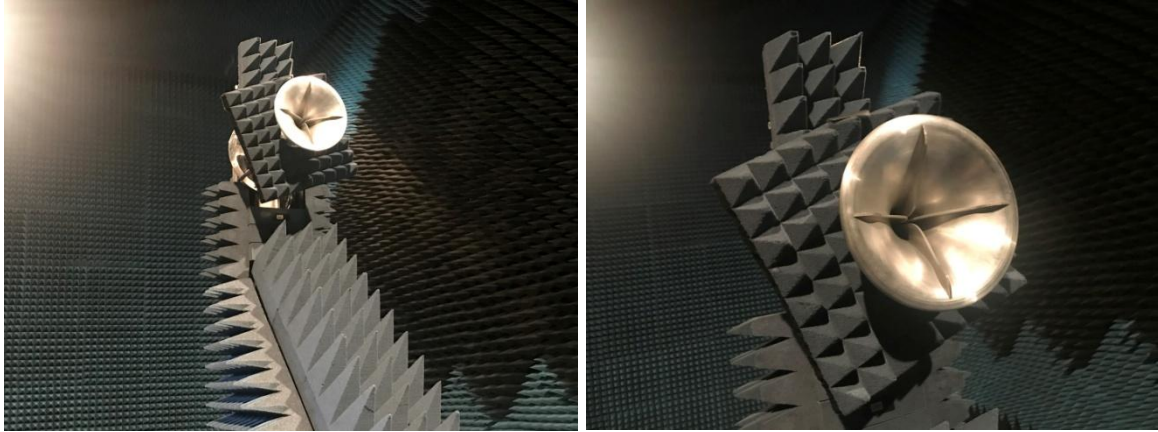


Fig 3.3. Measurement of the QRFH in the anechoic chamber of Yebes (CDTAC).

It is a broadband antenna, so it has been necessary to use several probes for the measurements, since each probe works in a specific frequency band. The different probes that were used, and the frequency bands measured with each one, are shown in Table II.

Probe	Frequency band	Delta	Frequency points
WR430	2-3 GHz	0.5 GHz	3
WR229	3.5-4.5 GHz	0.5 GHz	3
WR159	5-7.5 GHz	0.5 GHz	6
WR112	8-9.5 GHz	0.5 GHz	4
WR75	10-15 GHz	0.5 GHz	11

Table II. Antenna probes used for the radiation pattern measurements.

The measurement has been done with the Spherical Near Field System, with two linear polarizations (LIN-0 and LIN-90). Each antenna port has been measured separately, loading the other port with 50 Ω .

The complete set of measured radiation patterns of the antenna are plotted in the Appendix A. The cuts at 0, 45 and 90 degrees are shown, corresponding to the E, D and H planes respectively. The E-Plane is the one aligned with the port.

3.3 Phase center

The location of the phase centre, as a function of frequency, obtained from the measured radiation patterns is plotted in Fig 3.4. This position has been defined with respect to the aperture, in such a way that negative values correspond to positions inside the horn.

The optimal position of the feed with respect the focus of the radio telescope has been calculated as the one that maximized the average aperture efficiency in the broad band. The obtained values were -52.99 mm and -53.19 mm for port B and A respectively, which means that the optimal for this antenna is about -53 mm from the aperture (the tolerances in the measurement are in the order of 1 mm).

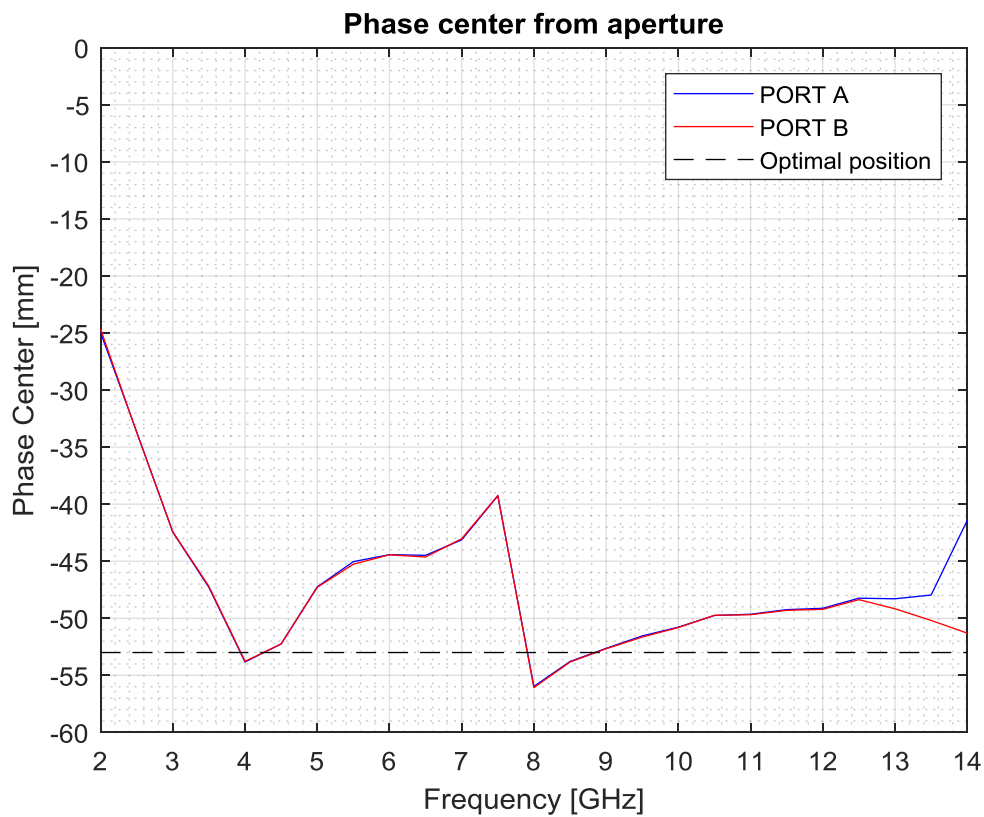


Fig 3.4. Measured phase center position with respect to the aperture.

4 Aperture efficiency

The aperture efficiency of the VGOS 13.2m reflector antenna illuminated by the QRFH antenna can be calculated using the measured radiation patterns of the feed. The parameters of the VGOS antenna used for the calculation are the following:

- Antenna type: Ring Focus [6][7]
- Subtended half-angle: 65°
- Diameter of the parabola: 13.2 m
- Focal length: 3.7 m
- Subreflector diameter: 1.48 m

The aperture efficiency has been calculated analytically, with the expressions in [8]-[11]. The results obtained from the analytical expressions have also been reproduced with optical physics, using GRASP, without significant differences. The overall aperture efficiency of each antenna port is shown in Fig 4.1. In Fig 4.2 the aperture efficiency is decomposed by a set of subefficiencies: illumination, spill-over, polarization and phase.

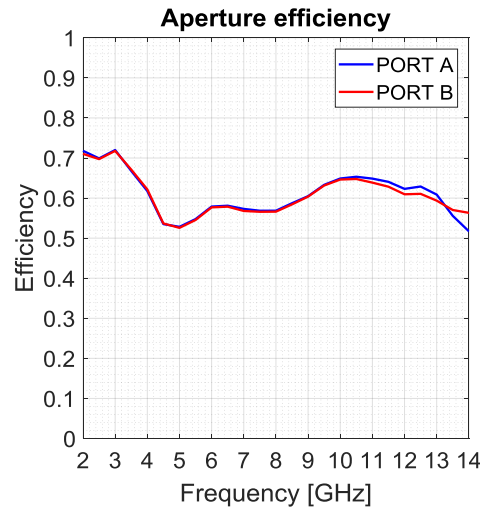


Fig 4.1. Estimated aperture efficiency of QRFH for each port.

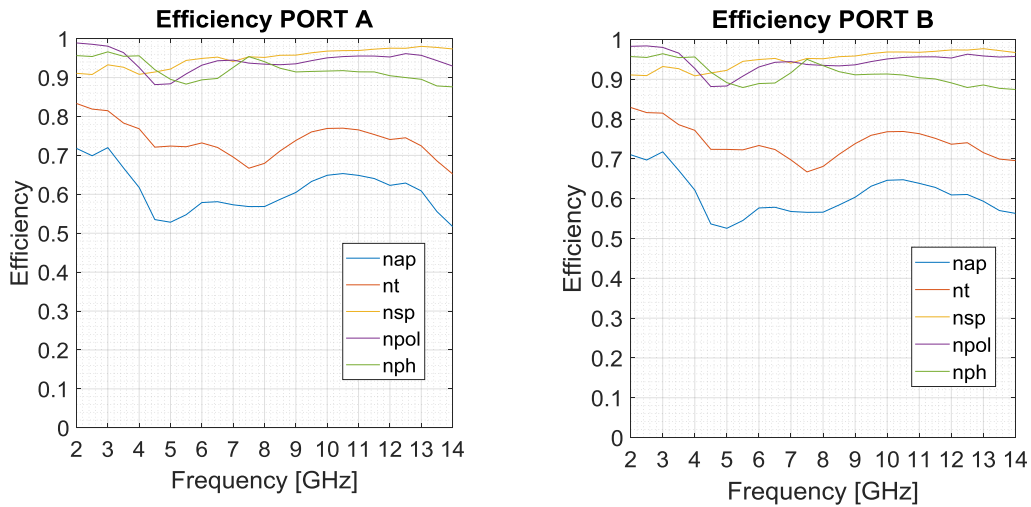


Fig 4.2. Estimated subefficiencies of QRFH for each port.

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Appendix A. Radiation patterns

